



Upper Boundary Extension of the NASA Ames Mars General Circulation Model

A. S. Brecht¹ (amanda.s.brecht@nasa.gov), J. L. Hollingsworth¹, M. A. Kahre¹, J. R. Schaeffer¹
¹ NASA Ames Research Center, Moffett Field, CA USA

Introduction

- Extending the NASA Ames Mars General Circulation Model (MGCM) upper boundary will expand our understanding of the connection between the lower and upper atmosphere of Mars through the middle atmosphere.
- The extension's main requirements is incorporation of Non-local thermodynamic equilibrium (NLTE) heating (visible) and cooling (infrared).
- NLTE occurs when energy is exchanged more rapidly with the radiation field (or other energy sources) rather than collisions with other molecules.
- Without NLTE, above ~80km/~60km in Mars' atmosphere the IR/visible heating rates are overestimated.
- Currently NLTE has been applied successfully into the 1D RT code and is in progress for the 3D application.

1D Radiative Transfer Code

- RT code is based on two-stream approximation; quadrature in the visible, hemispheric mean in the IR
- Calculates fluxes at layer boundaries and at the surface; heating rates are calculated from these fluxes
- Includes CO₂ and H₂O gas opacity; aerosol optical properties for dust and water ice clouds
- Correlated-k gas opacities; calculated off-line on pressure and temperature grid
- Pressure ranges: 10⁻⁶ to 10⁻¹⁴ mbar (pressure at each grid point 10x larger than previous)
- Temperature ranges: 50 to 350 K, ΔT = 50
- 12 spectral intervals (7 visible, 8 IR) with 17 Gauss points per interval
- IR spectral range: 1000–4.5 μm; Visible spectral range: 4.5–0.24 μm
- For more information about the RT code go to: <http://space-science.arc.nasa.gov/mars-climate-modeling-group/>

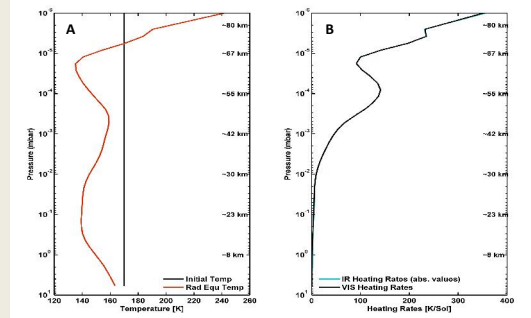
INITIAL CONDITIONS:

- Isothermal; 170 K
- Not connected to 3D model for initial testing
- P_{top} = 2E-6 mbar; 40 layers
- Surface pressure = 7 mbar; Albedo = 0.24, Dust free
- SZA = 80 degrees (for radiative equilibrium case)

Strategy for 1D RT code

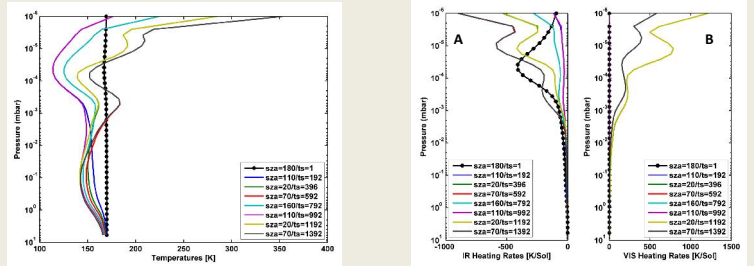
- Extend tropopause pressure from 8E-4 mbar to 2E-6 mbar (to enter NLTE region)
- Increase sigma layers from 24 to 40 (provide more resolution)
- Implement NLTE additions
 - Solar heating rate corrections from Table 1 in López-Valverde et al. (1998)
 - CO₂ 15 μm cooling parameterization adapted from Bougher et al. (2006)
- Increase IR spectral bands from 5 bands to 8 bands
- Check model to see if it goes to equilibrium and if it is stable

Radiative Equilibrium Temperature and Heating Rates



- Panel A represents a radiative equilibrium temperature profile (red line)
- The black line is the initial temperature profile.
- Panel B is the corresponding IR and Visible heating rates.

Temperature and Heating Rates: Time-Evolving



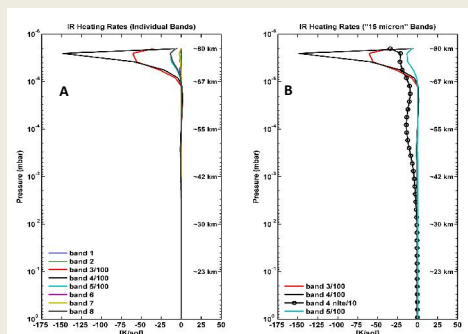
- Series of temperature profiles
- The RT code was time-evolved by cycling through SZA
- Each line represents a SZA at a specific timestep (ts)

- Panel A is the corresponding time evolving IR heating rates and panel B is the corresponding Visible heating rates

IR Spectral Bands

Band	Wavelength (μm)	
1	1000.0–33.3	
2	33.3–18.2	
3	18.2–15.7	Wing of 15 micron CO ₂ band
4	15.7–14.2	Core of 15 micron CO ₂ band
5	14.2–12.9	Wing of 15 micron CO ₂ band
6	12.9–10.3	12 micron water ice band
7	10.3–7.8	9 micron dust band
8	7.8–4.5	

Radiative Equilibrium IR Heat Rates



- All LTE band are represented in panel A (bands 3-5 are scaled by 100)
- Bands 3-5 are replaced with the corrected NLTE 15 micron heating rates (these bands are shown in panel B)

Conclusion/Future Work

- Adding NLTE provides smaller heating rates; which is the expected result
- 1D RT code does come to equilibrium and it is stable when evolved in time
- More analysis is needed for the upper boundary
- These 1D RT code updates will be implemented into the 3D code
- Analysis of MGCM results with the vertical extension and NLTE are still in progress

- Sample of 3D IR heating rates
- 5 spectral bands (not comparable to other results on poster)
- Microphysics off
- Constant dust everywhere with a visible dust optical depth = 0.3 and Conrath-ν = 0.03
- Cold start; 100 sol runs (Ls=50 to Ls=90)
- Output: zonally averaged over 10 sols

